

## WHAT HAVE GALAXIES DONE WITH THEIR METALS?

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The eventual fate of a galaxy's metals is a direct tracer of its history of star formation, gas flows, and feedback processes. I will show that a simple model combining empirical star formation histories with the local relation between stellar mass, gas-phase metallicities, and star formation rates reproduces the metal distributions of  $z=0$  galaxies remarkably well. I will then present an accounting of metals made by  $\sim L^*$  galaxies at  $z=0$ , showing that the bulk of metals released by supernovae and AGB stars are no longer in galaxies, and are instead in the circumgalactic medium (CGM) or intergalactic medium. The COS-Halos survey has created a statistically-sampled map of the gaseous CGM of low redshift  $\sim L^*$  galaxies out to impact parameters  $\sim 150$  kpc using the Cosmic Origins Spectrograph (COS) on the Hubble Space Telescope. Using this map, I will show that the masses of metals found in the cool ( $10^4 < T < 10^5$  K) photoionized and the more highly ionized OVI-traced CGM could potentially account for all of the metals expelled from  $\sim L^*$  galaxies at  $z \sim 0$ .